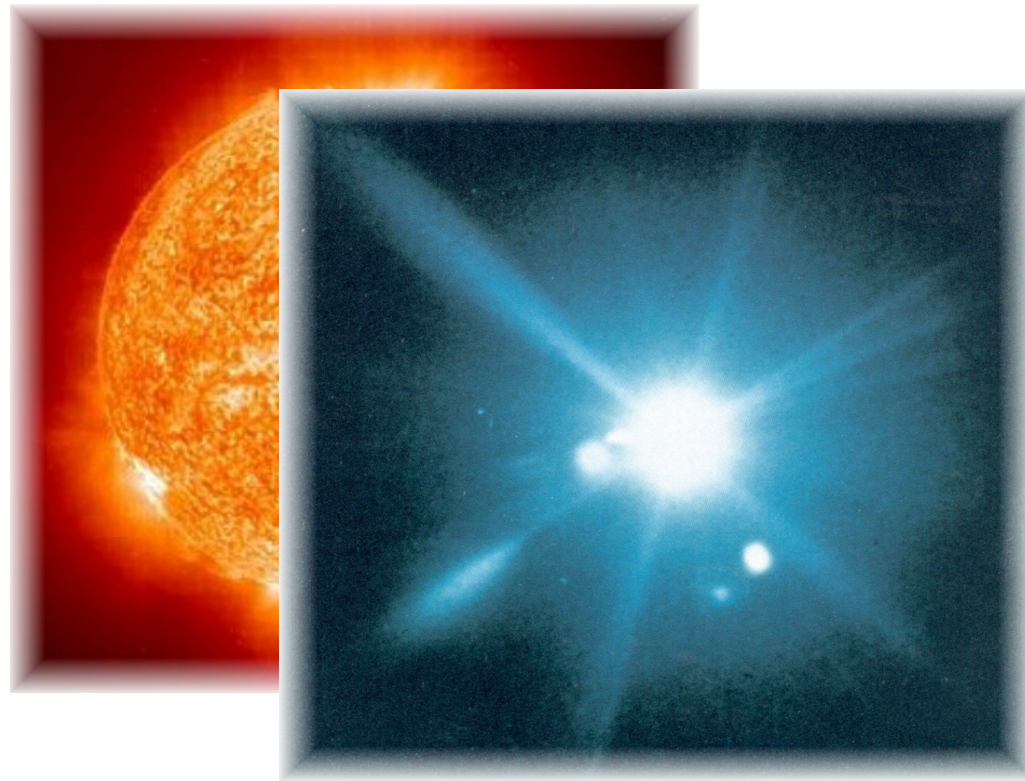


Introduction to synchrotron radiation

Martin Müller

Institut für Experimentelle und Angewandte Physik
der Christian-Albrechts-Universität zu Kiel



Brighter than
the sun!

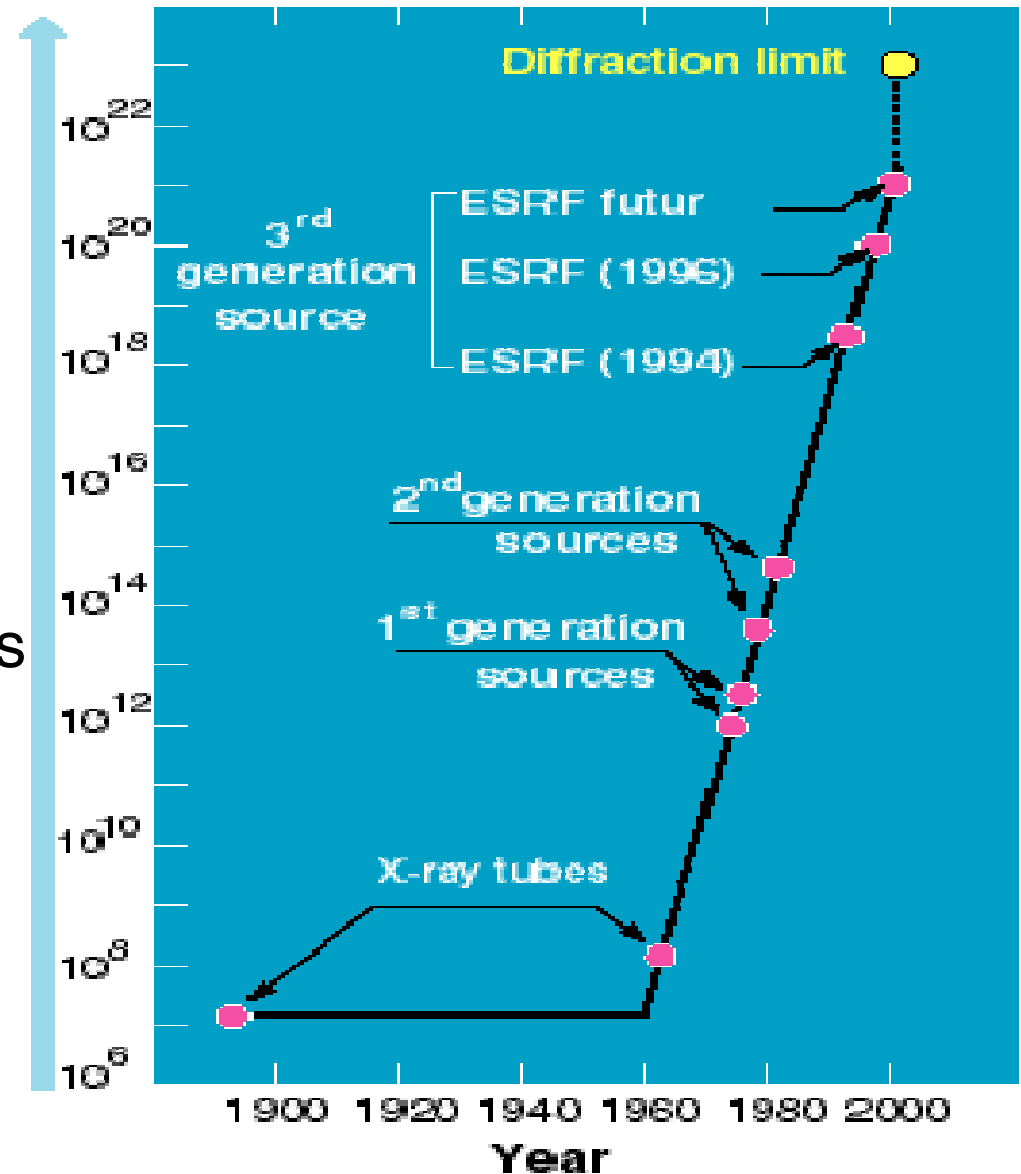
Synchrotronstrahlung

Brilliance of the X-ray beams
(photons / s / mm² / mrad² / 0.1% BW)

Synchrotron-Infos
auf Kieler Website:

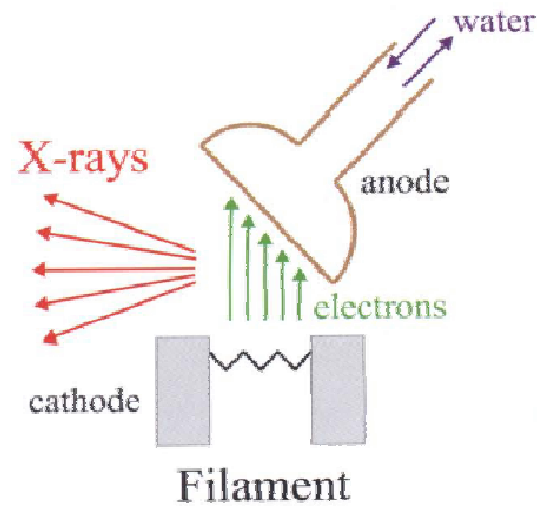
Komitee
Forschung mit
Synchrotronstrahlung

<http://www.physik.uni-kiel.de/kfs>

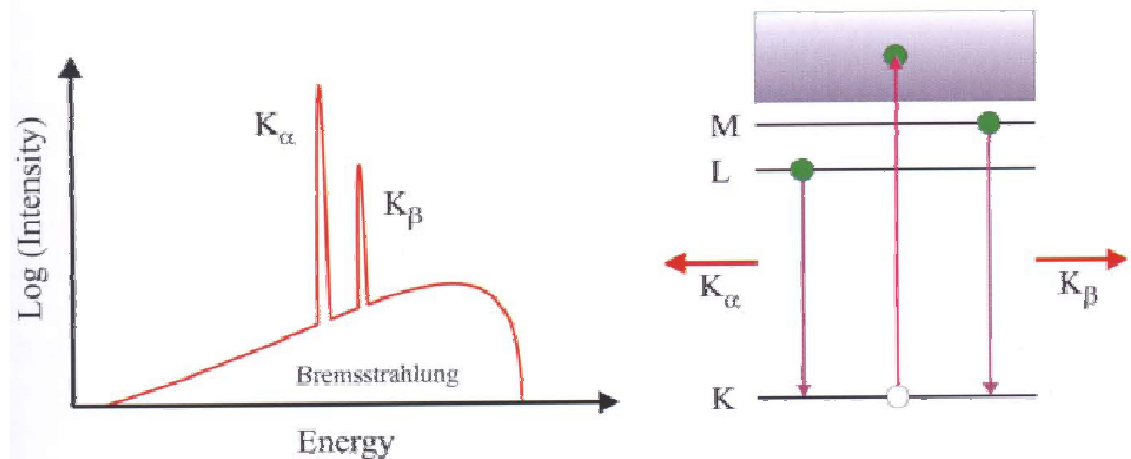
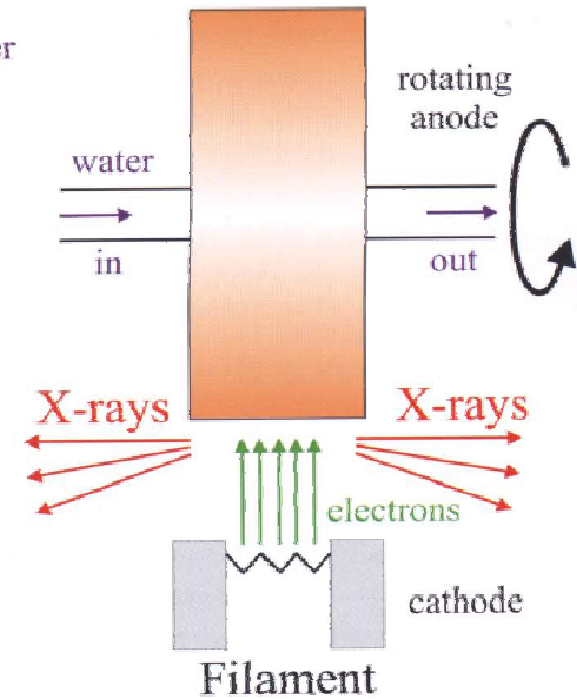


Konventionelle Röntgenquellen

Coolidge Tube



Rotating Anode



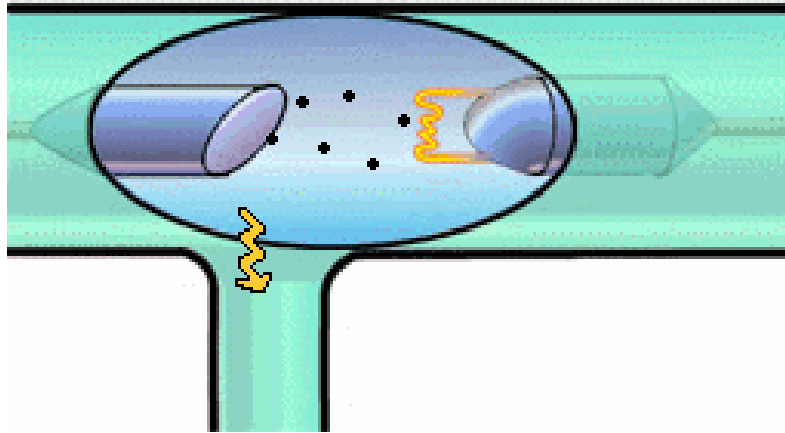
Als-Nielsen 2001, Abb. 2.1

Generation of X-rays...

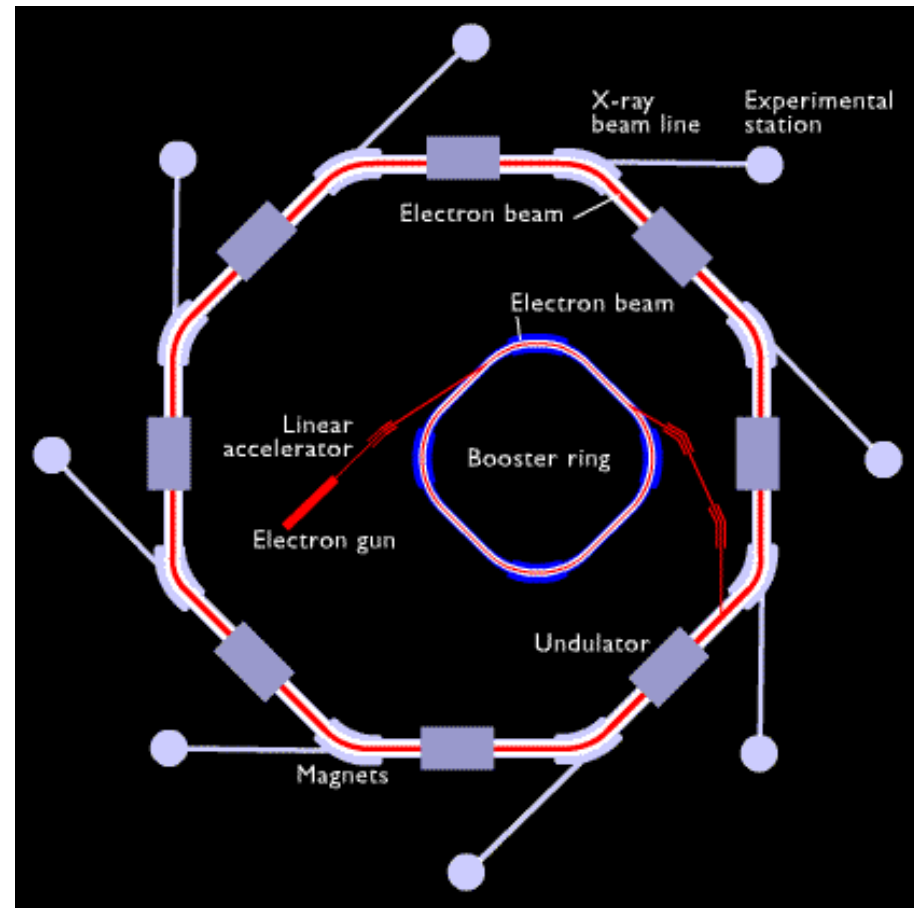
...by decelerated / accelerated electrons

X-ray tube (anode)

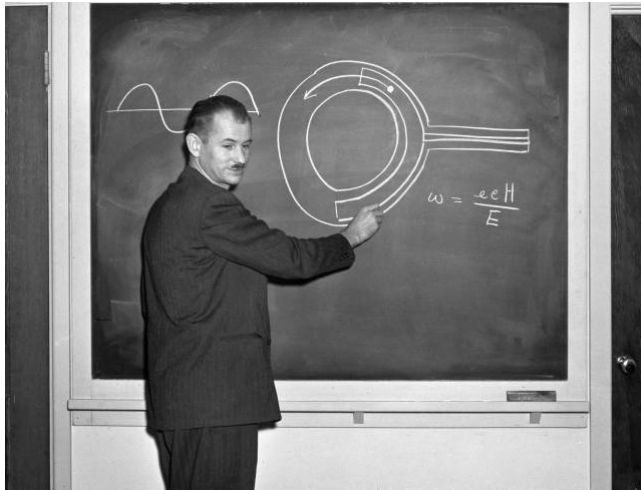
Electron Synchrotron



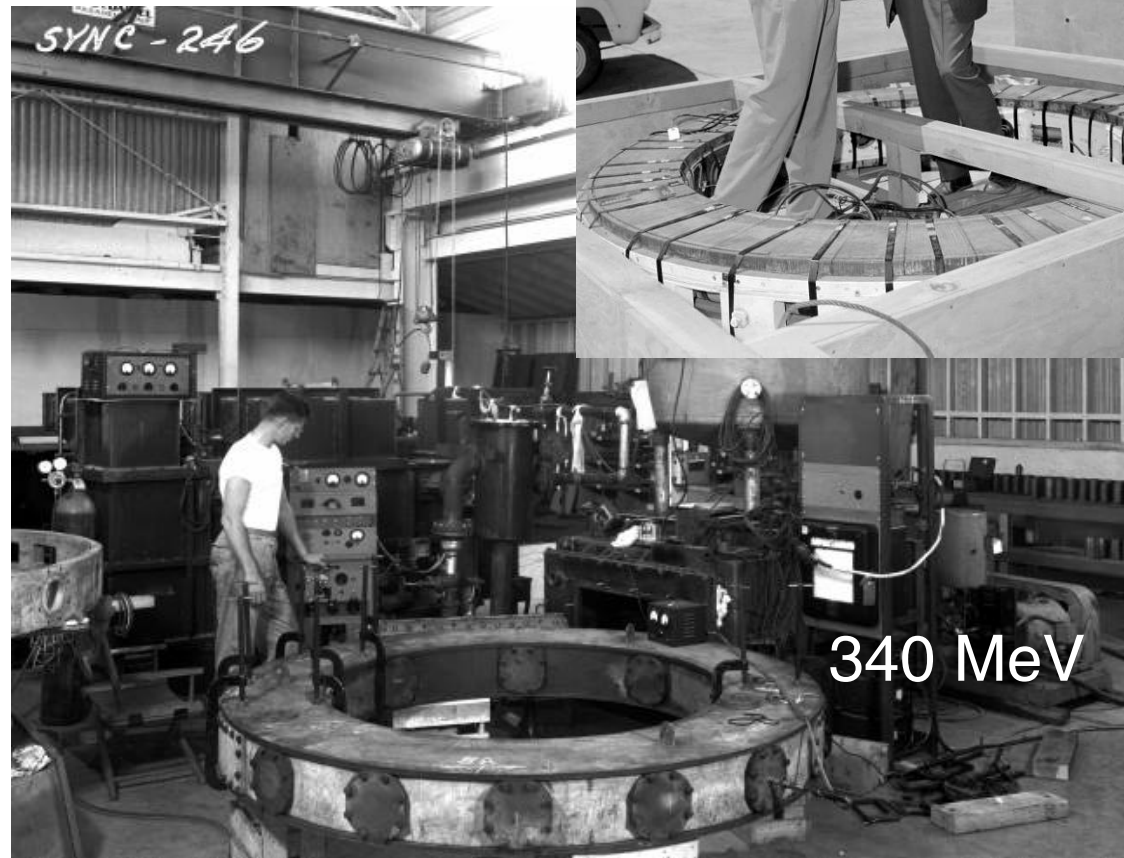
1895



The first synchrotron (1948)

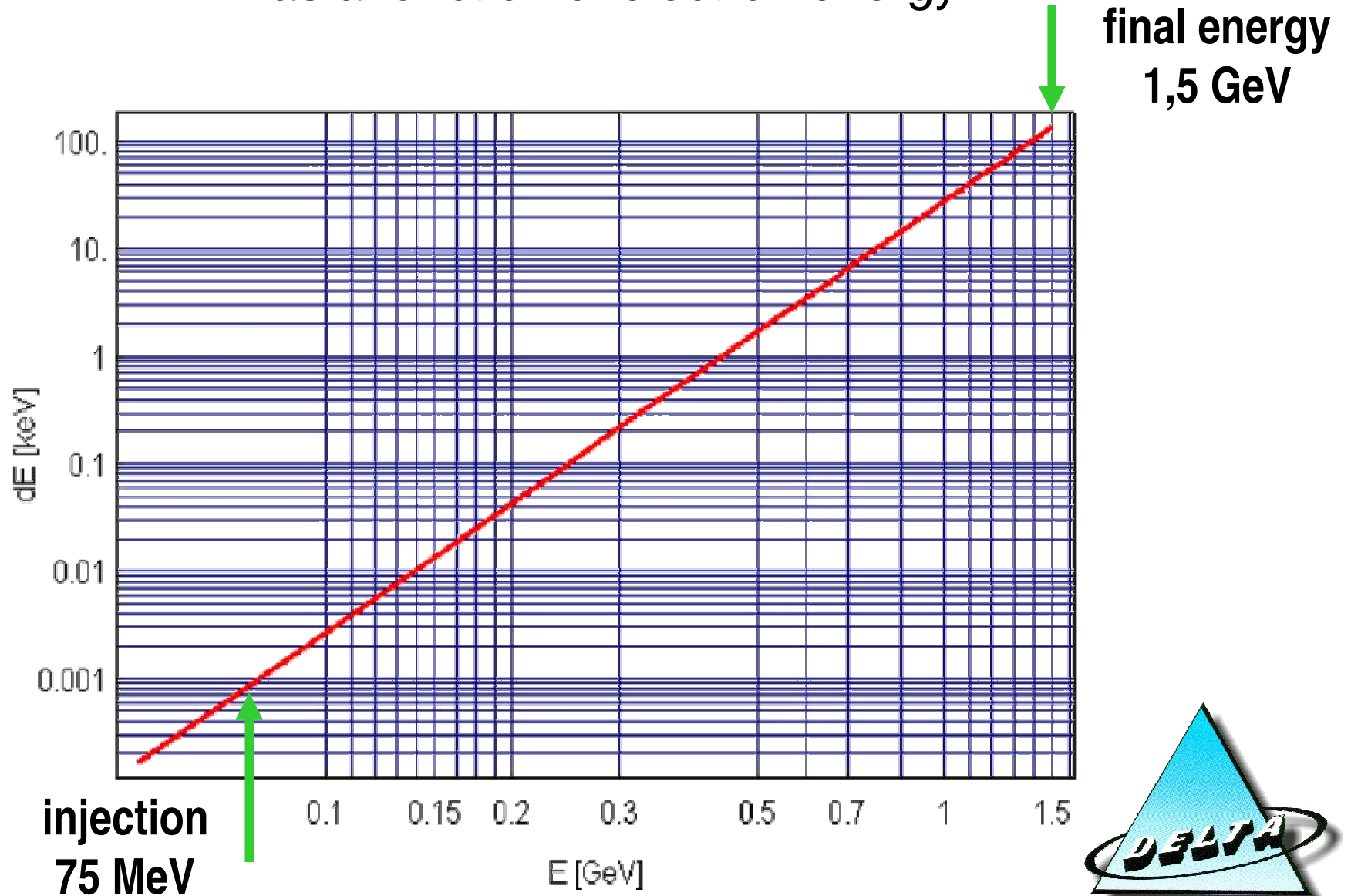


Edwin McMillan,
Berkeley Laboratory

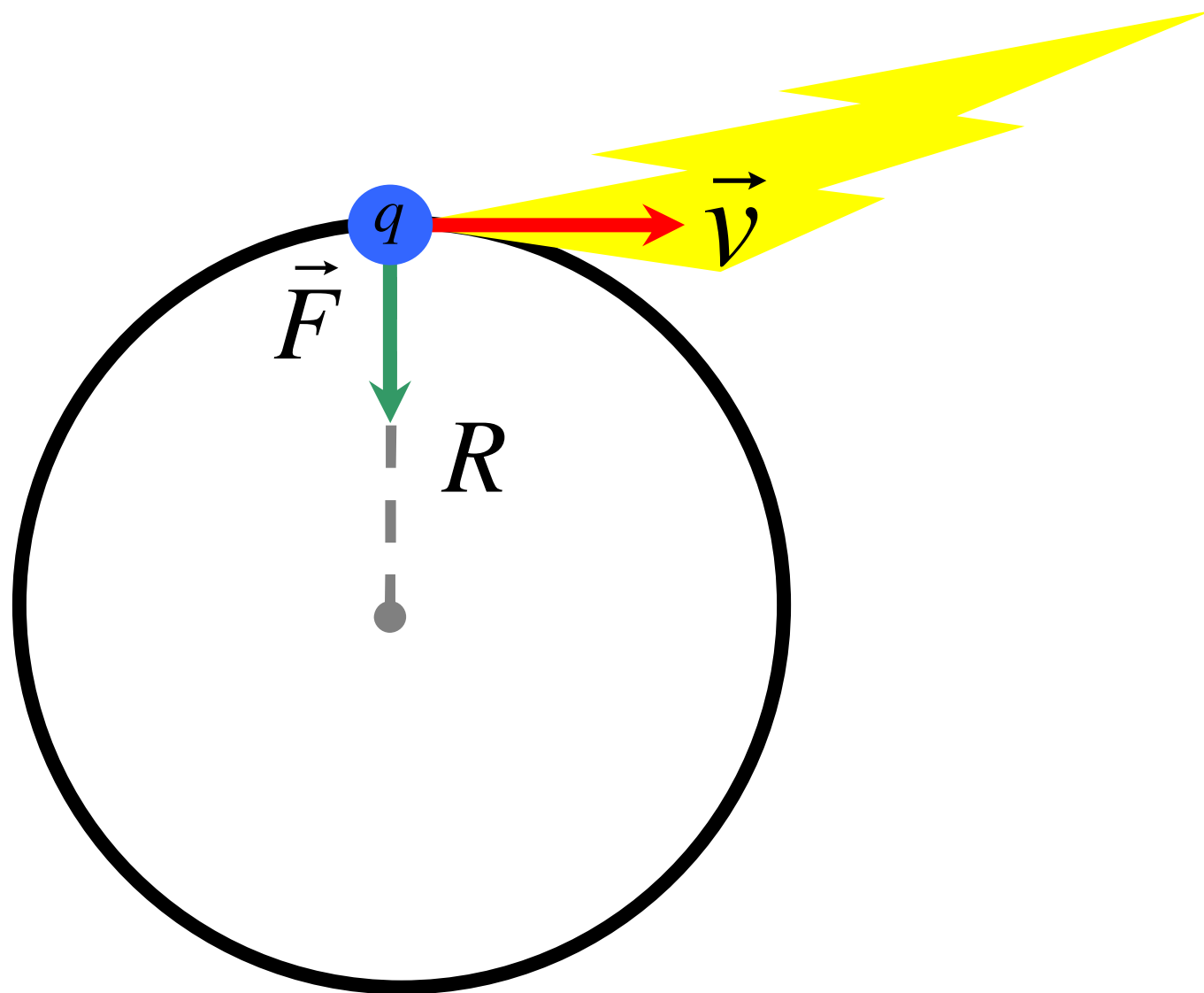


Energy loss per cycle

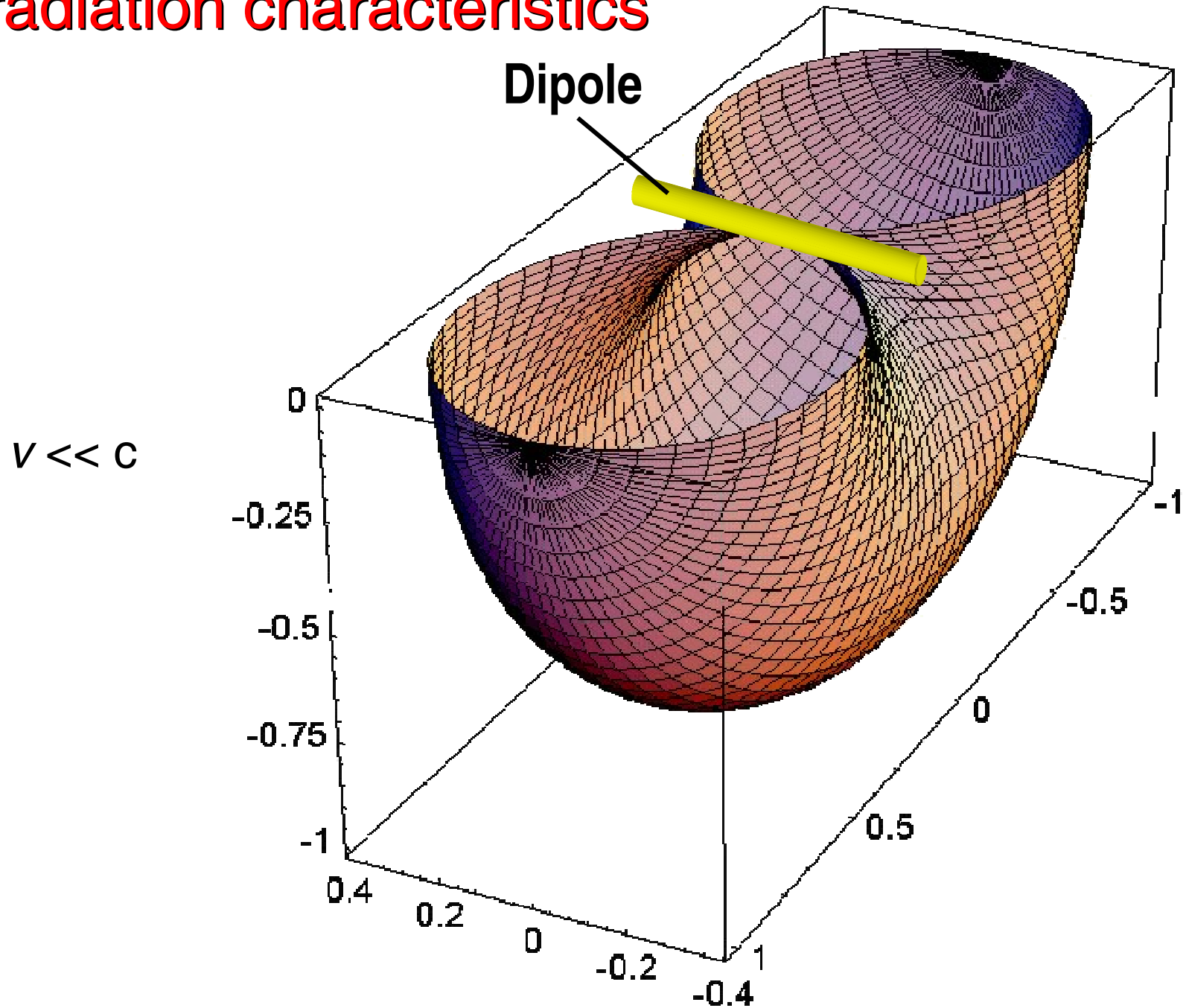
as a function of electron energy



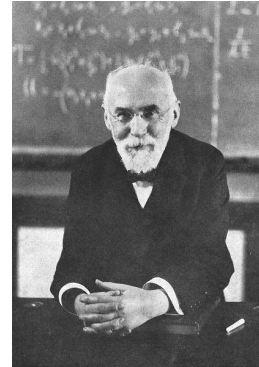
Charge (e.g. electron) in circular orbit



Dipole radiation characteristics



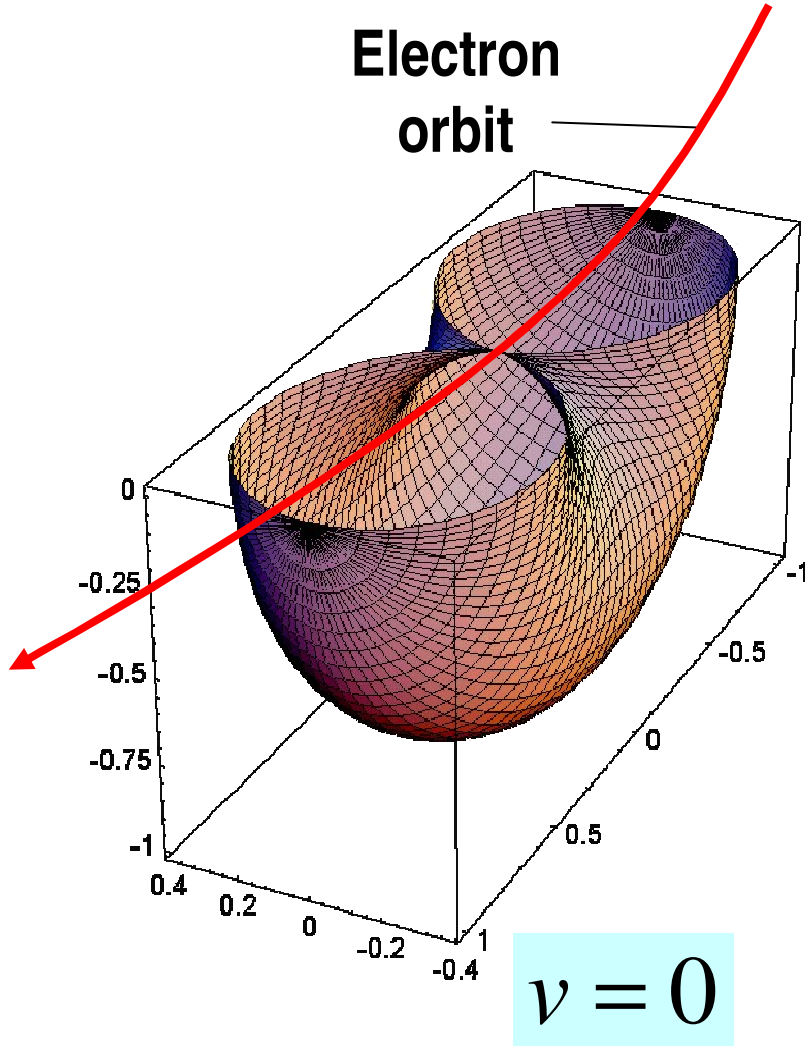
Lorentz transformation



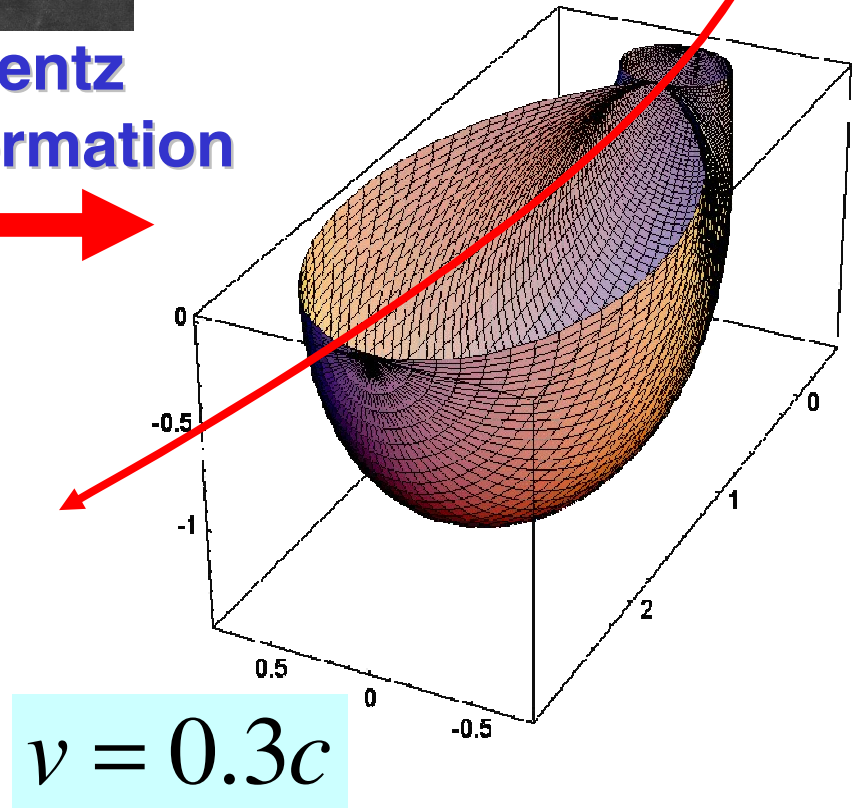
Lorentz transformation



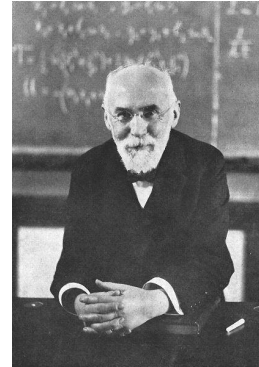
Electron orbit



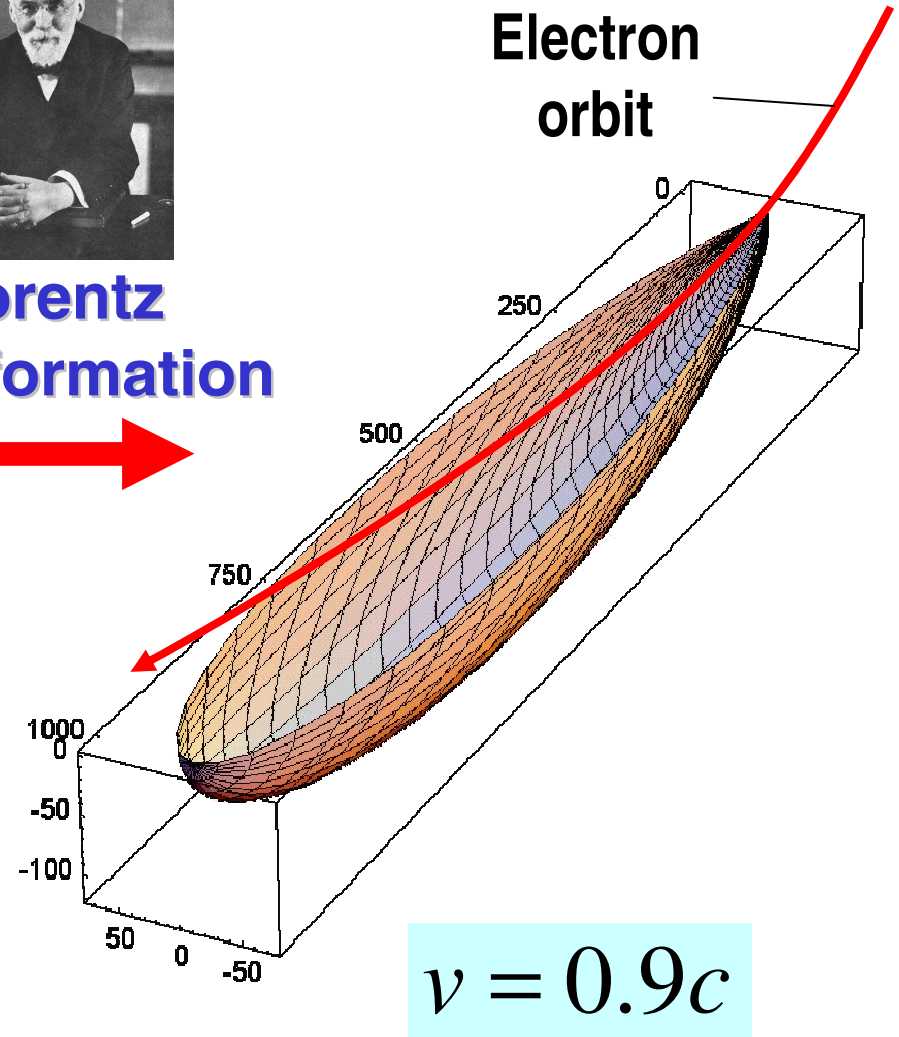
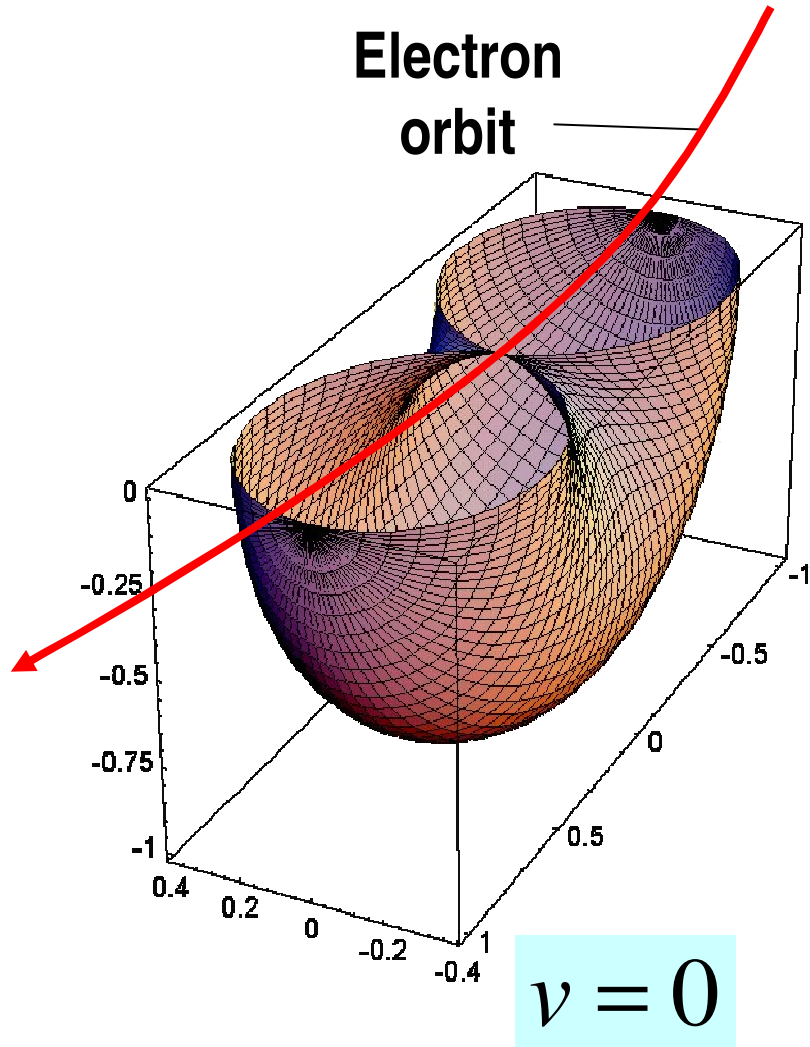
Electron orbit



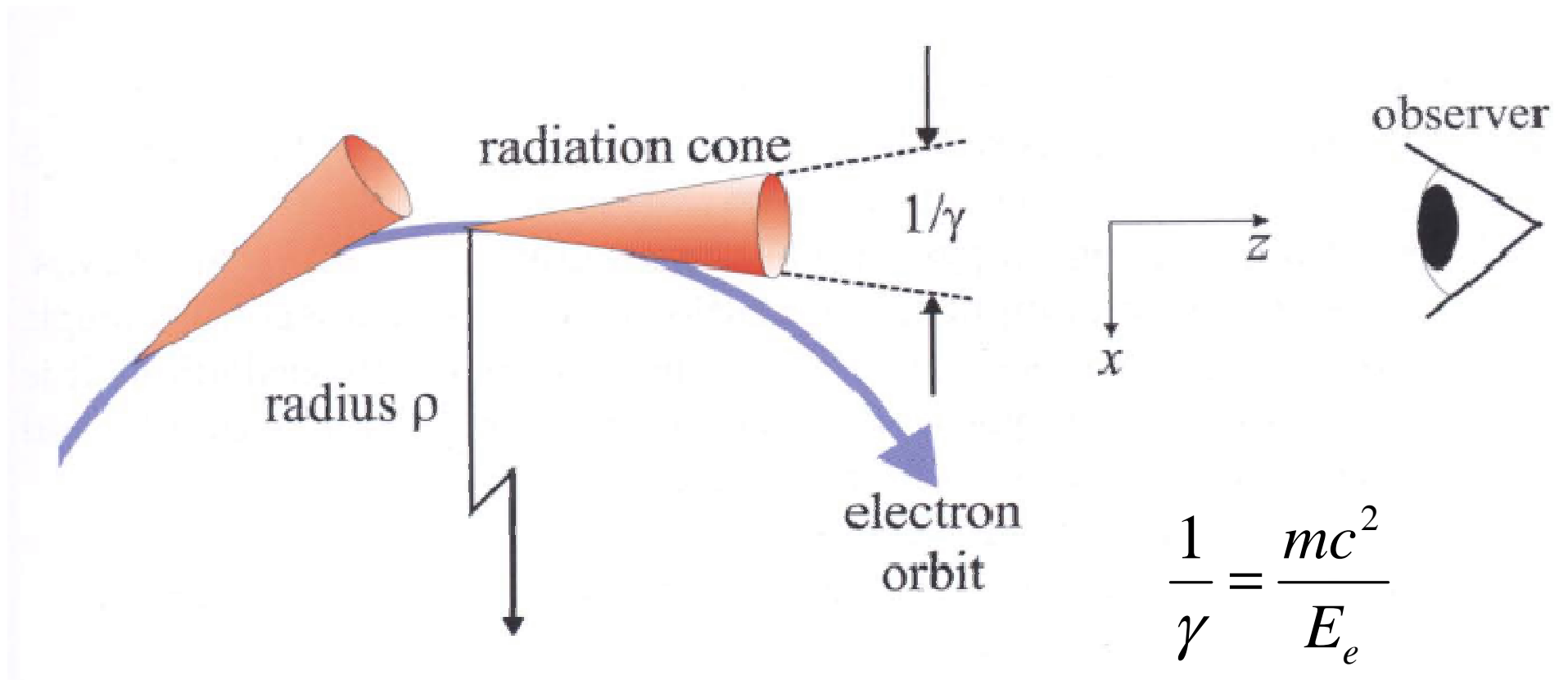
Lorentz transformation



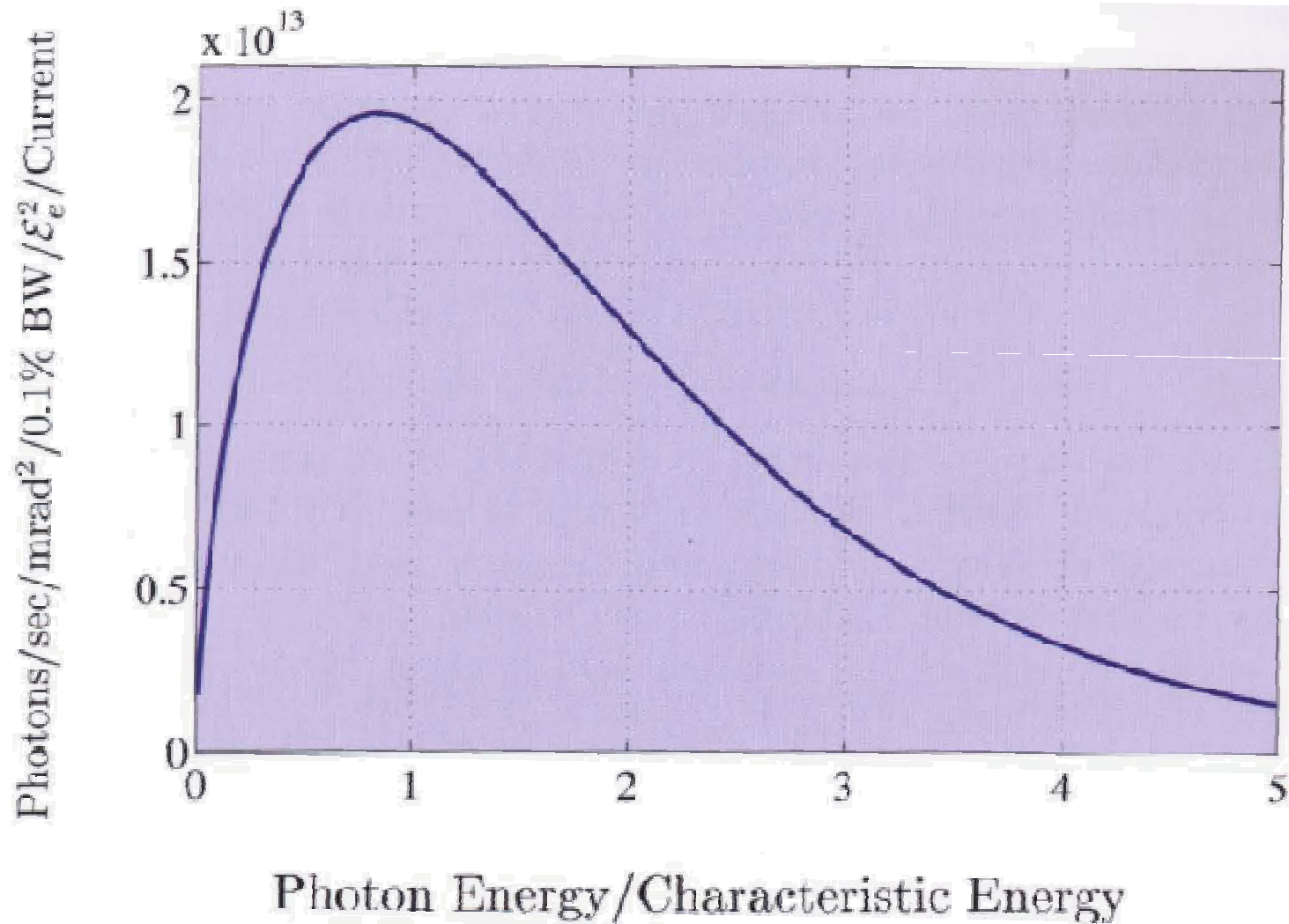
Lorentz transformation



Synchrotronstrahlung von einem Kreisbogen



Spektrum eines Ablenkmagneten

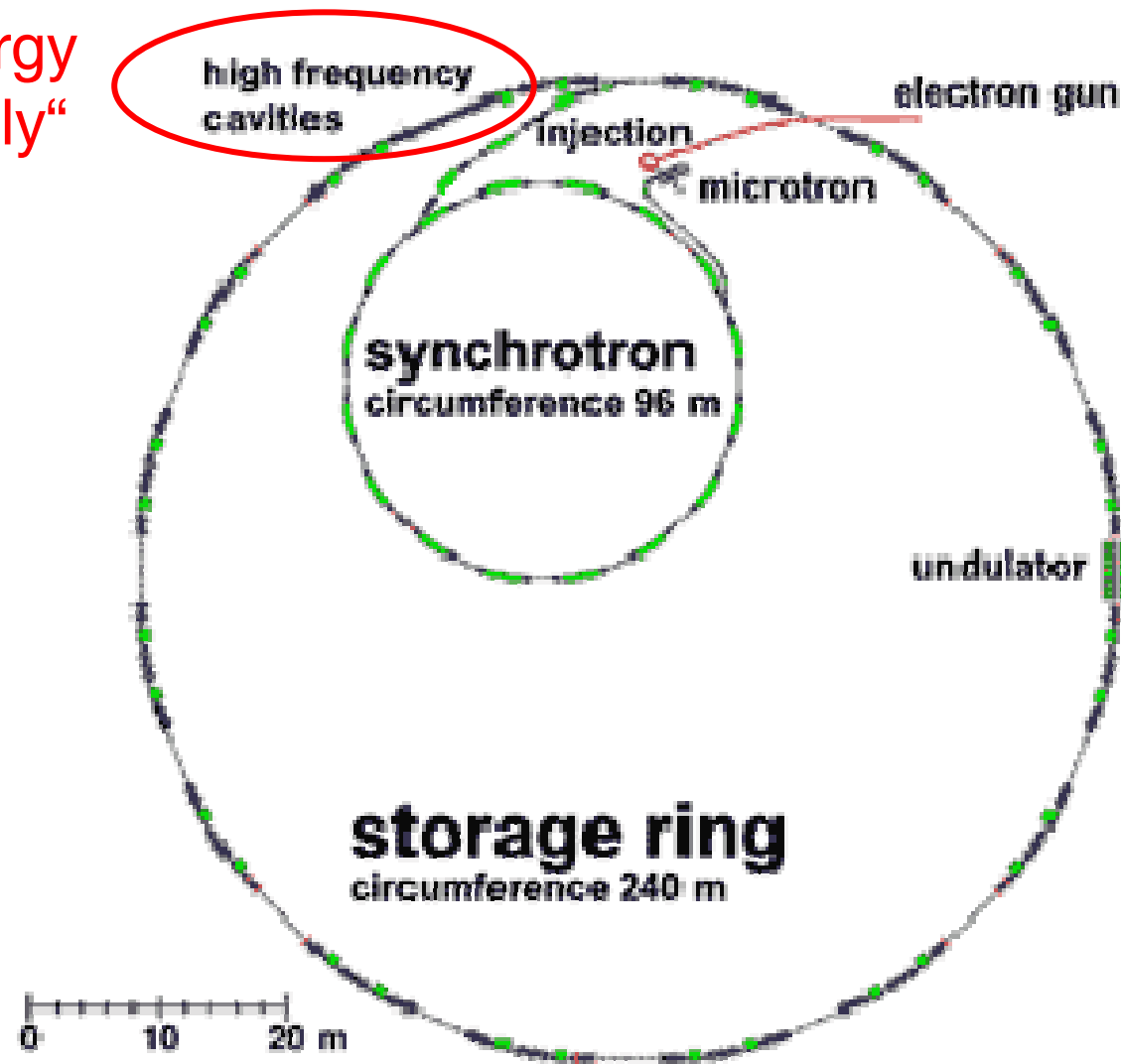


Beispiel: ESRF-Ablenkmagnet

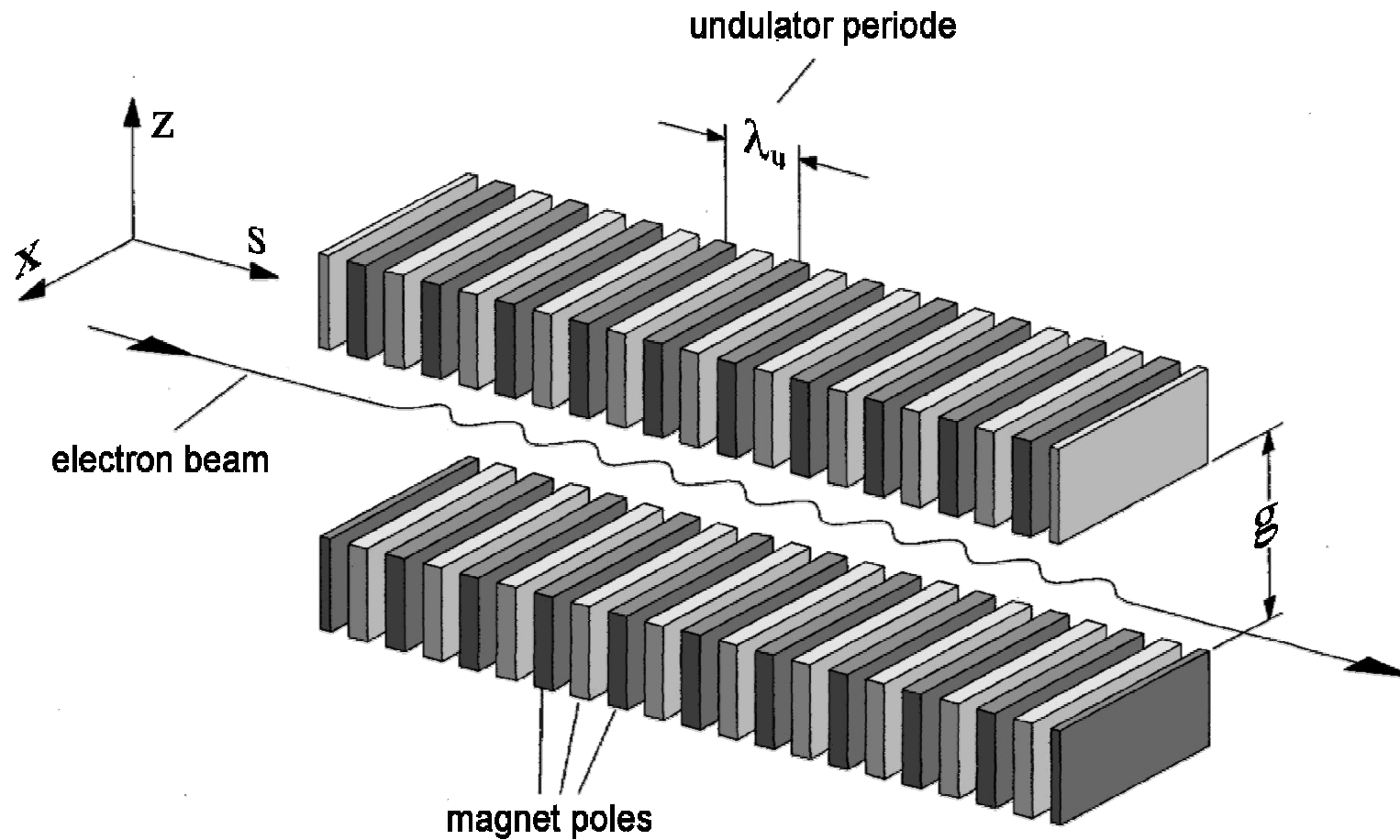
Als-Nielsen 2001, Abb. 2.5

3rd generation synchrotron (BESSY II)

„energy supply“

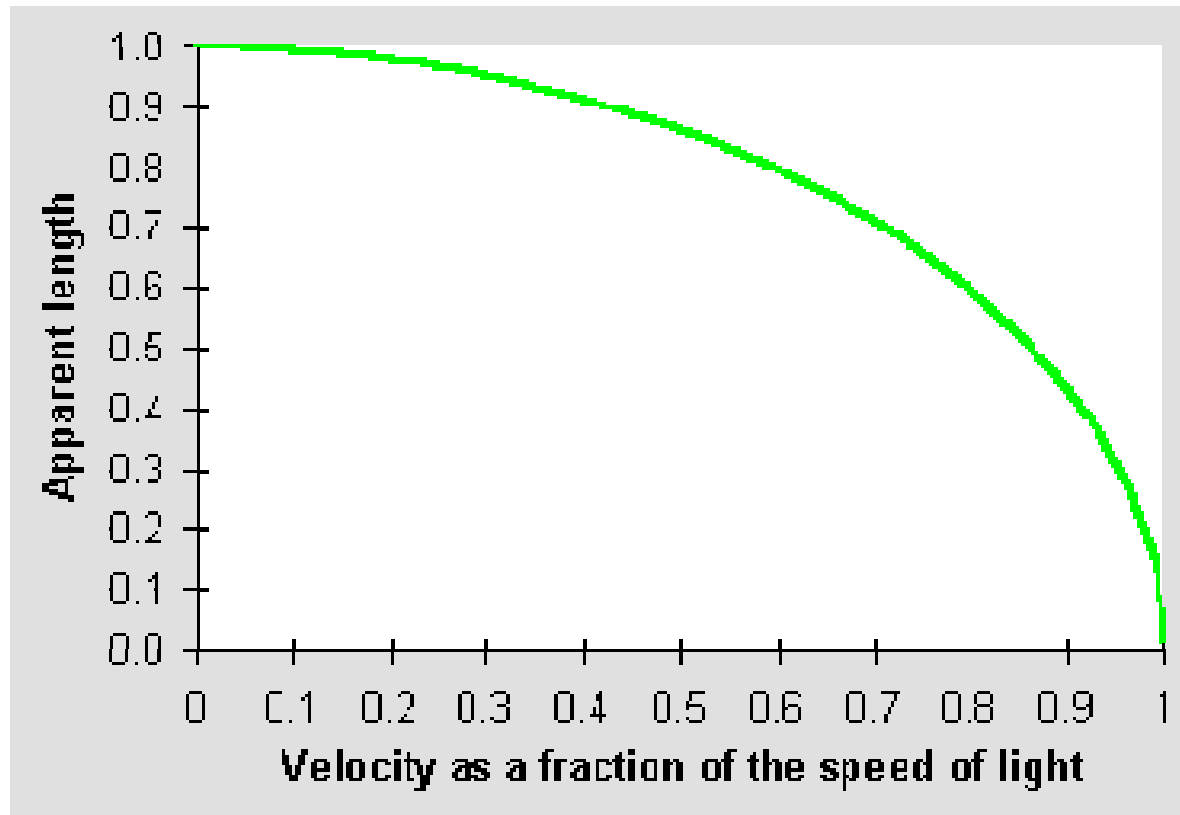


Principle of an undulator



constructive **interference** of all waves: **how?**

Lorentz contraction



$$l' = l \sqrt{1 - \frac{v^2}{c^2}} = l / \gamma$$

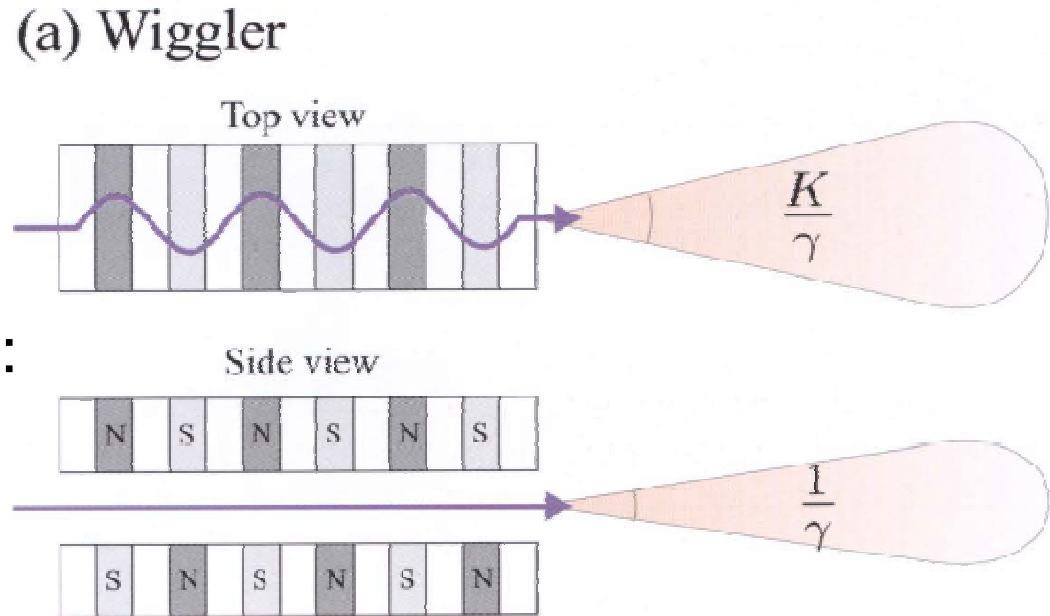
typical γ : $10^3 - 10^4$

additional factor = $1/(2\gamma)$ from Doppler effect

⇒ from **cm** (undulator structures) to **Å** (X-ray wavelengths)

Wiggler und Undulatoren

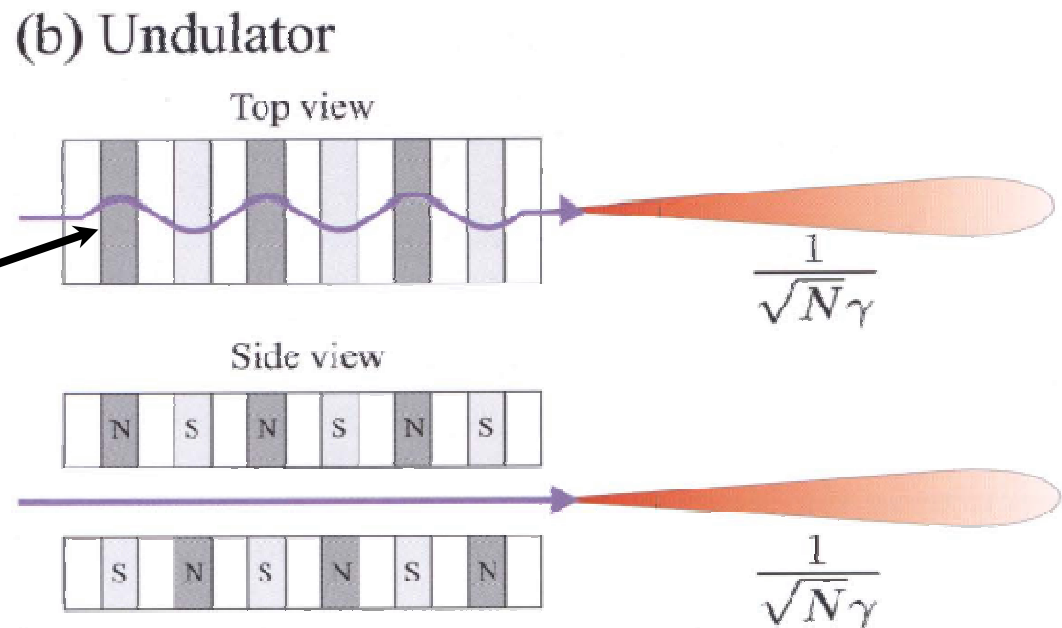
Summation von Intensitäten:
 $I \propto N$



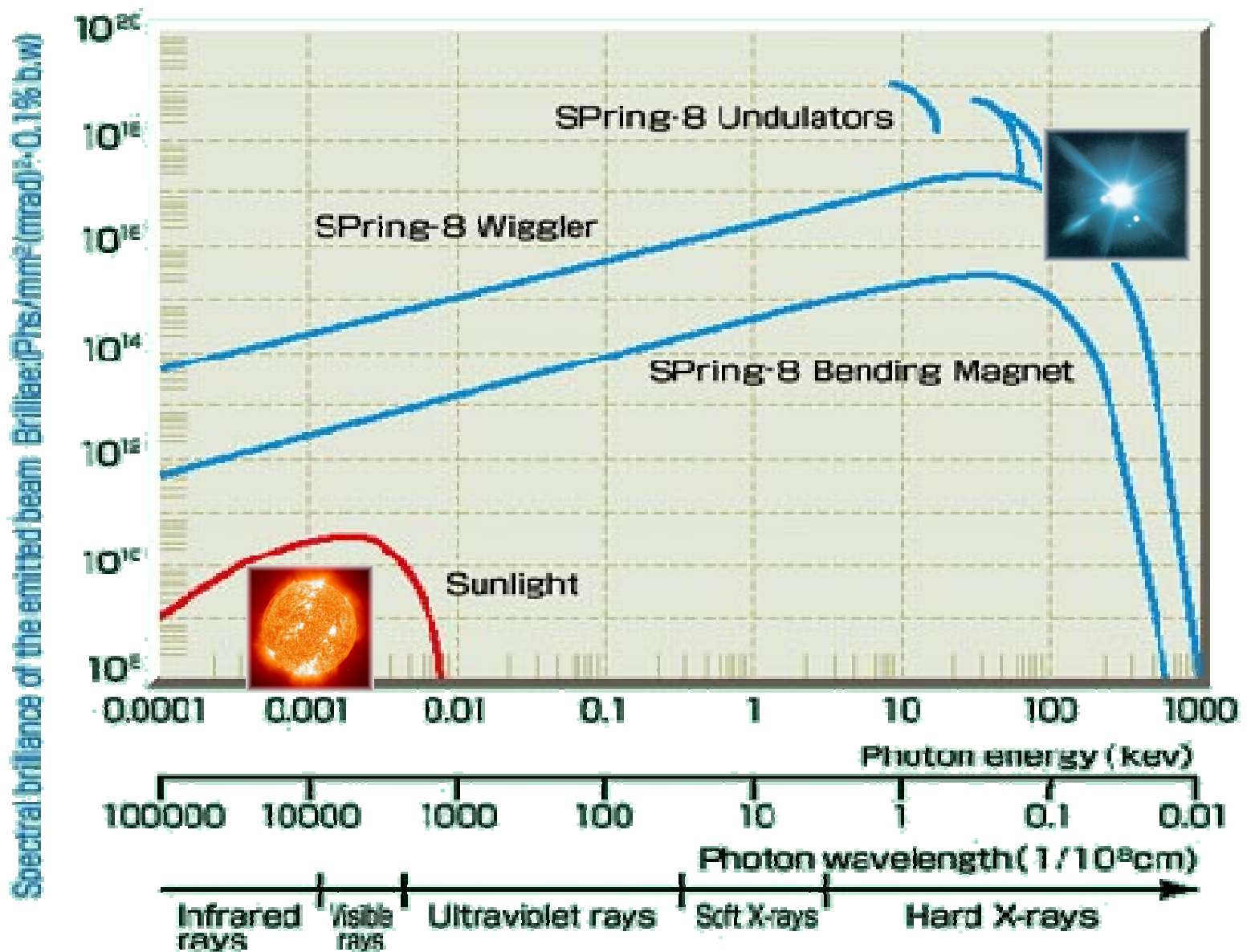
Als-Nielsen 2001, Abb. 2.7

Summation von Amplituden
(Oszillationen in Phase):
 $I \propto N^2$

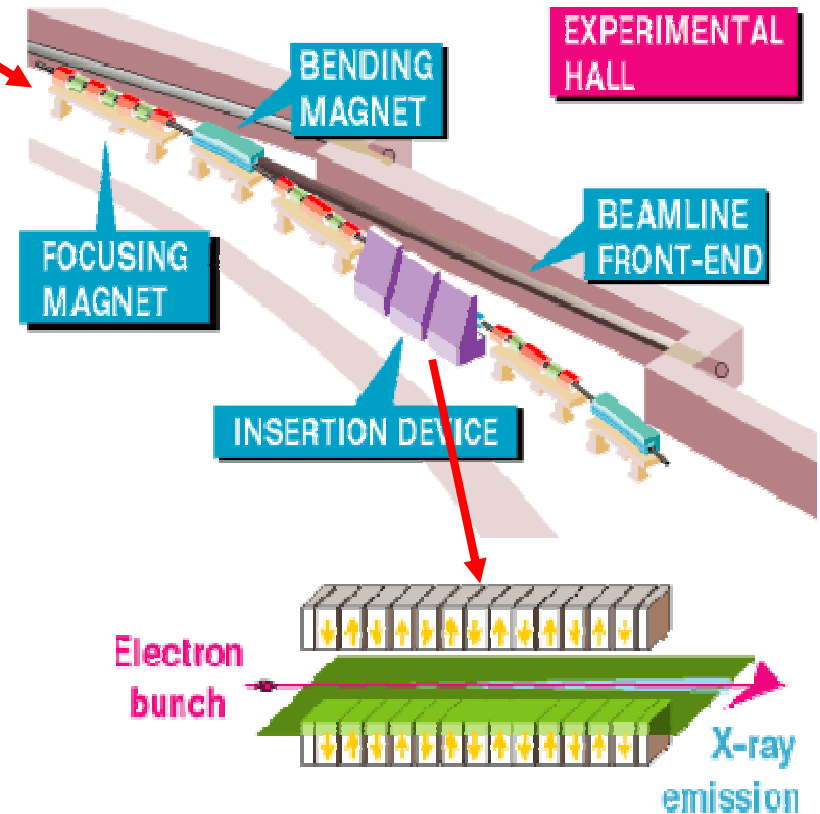
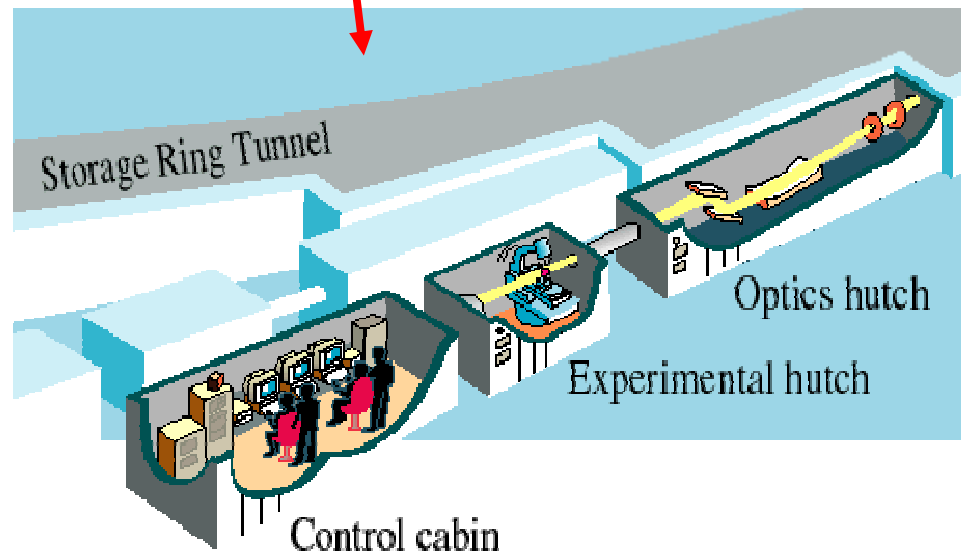
kleine Auslenkungen,
Winkel K



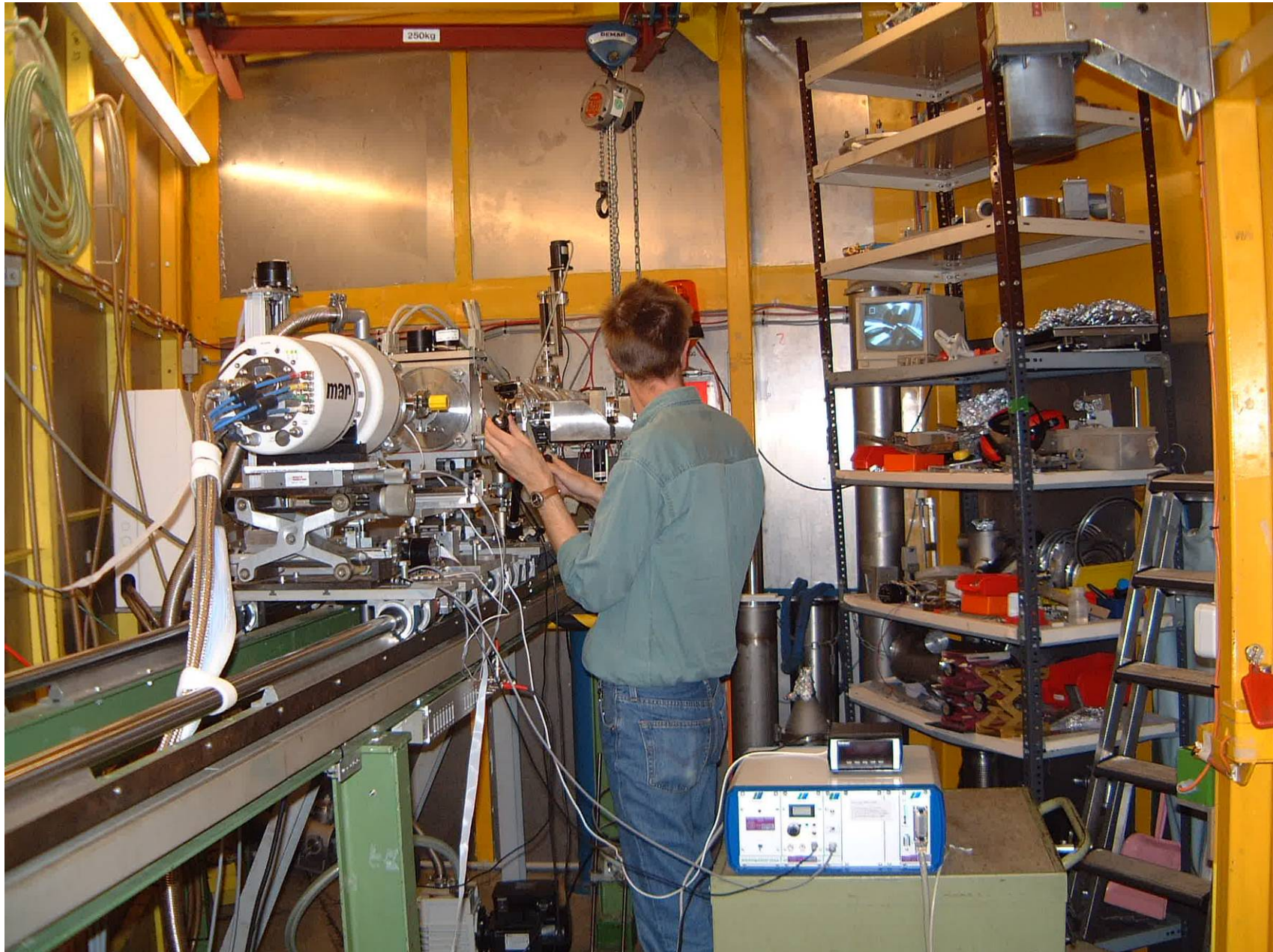
Spectral brilliance: Brighter than the sun!



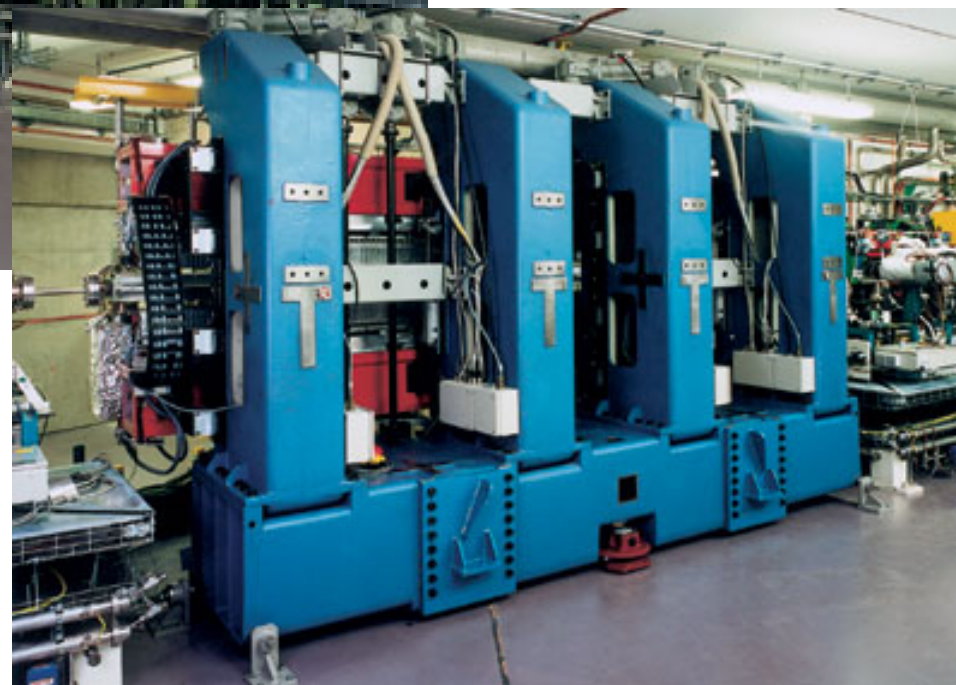
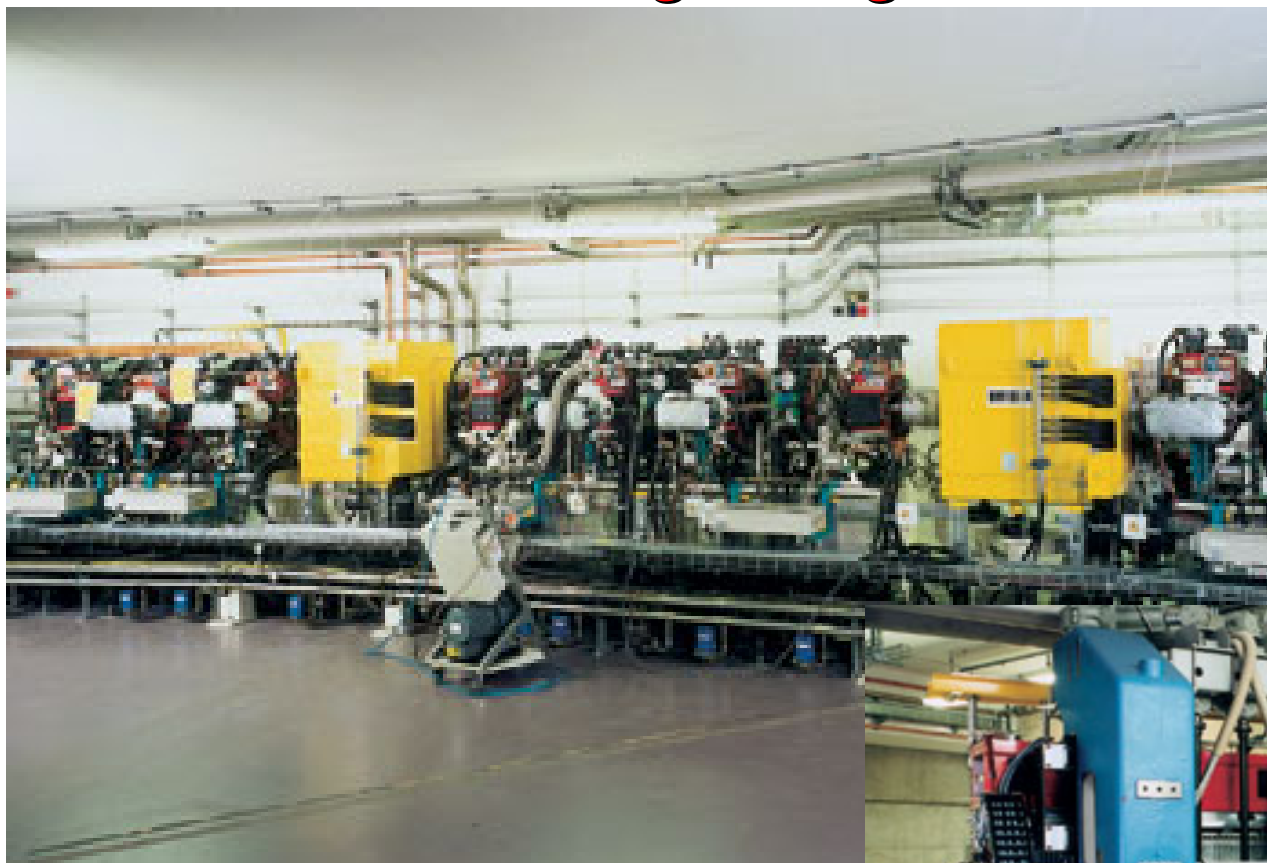
Synchrotrons today: user facilities



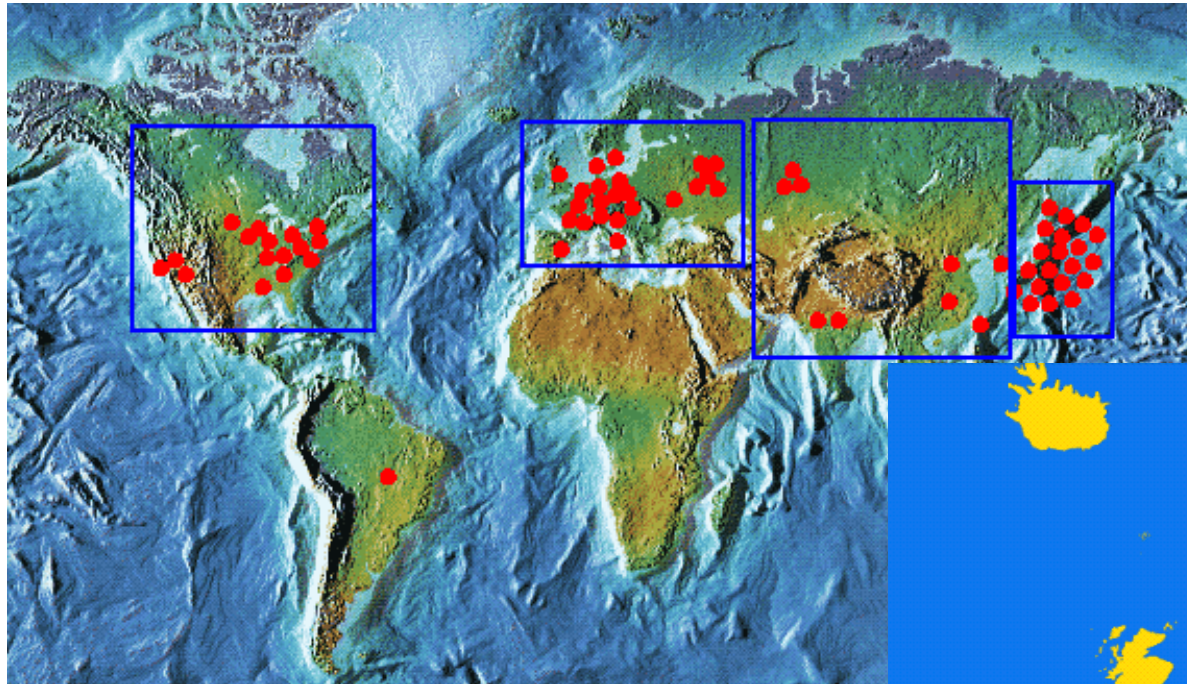
user@work



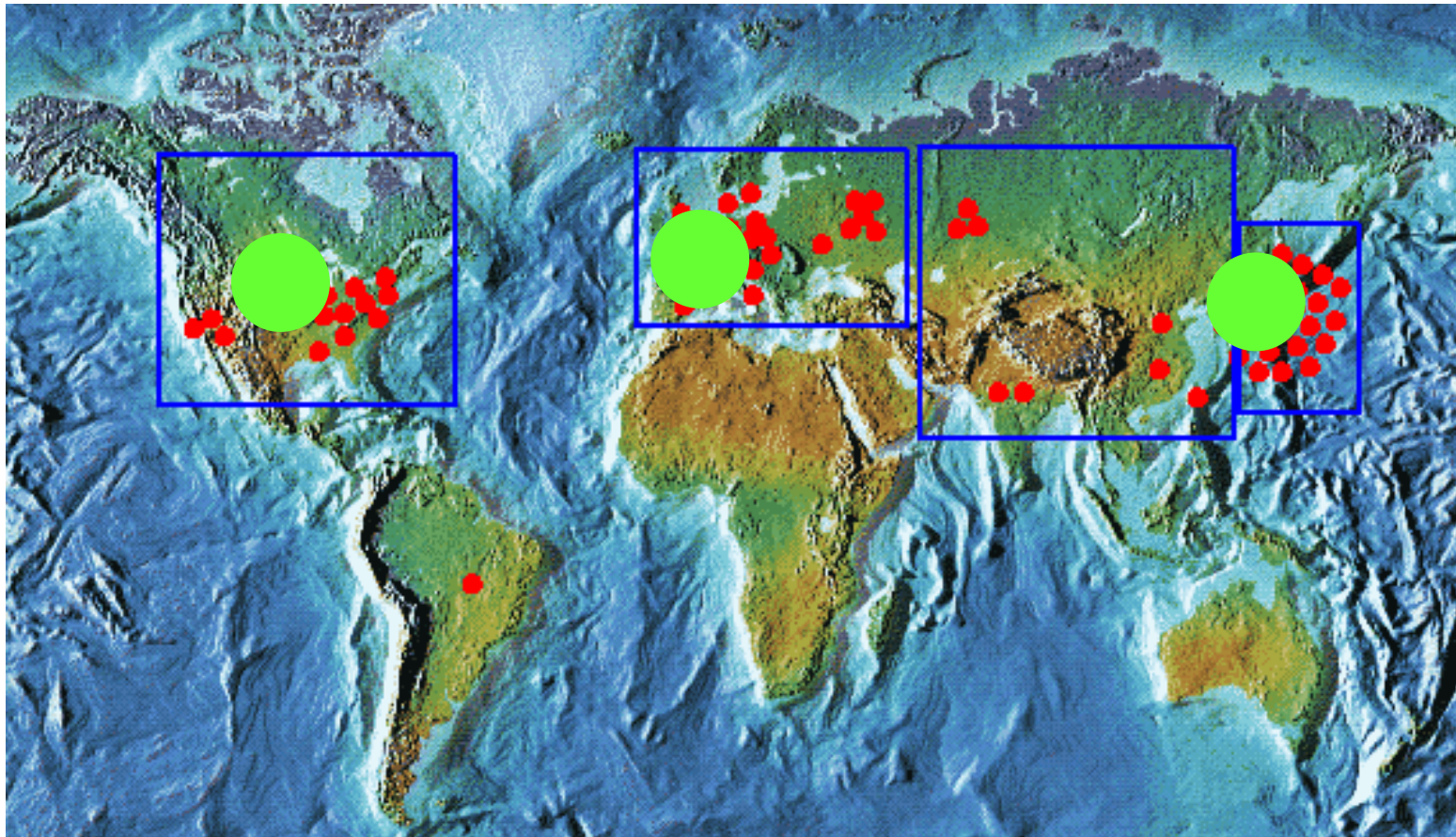
Storage ring and undulator



Synchrotrons worldwide



The three big synchrotrons





SPring-8: $E = 8 \text{ GeV}$, $U = 1436 \text{ m}$
Harima Science Garden City (Japan)



ADVANCED
PHOTON
SOURCE



APS: $E = 7 \text{ GeV}$, $U = 1104 \text{ m}$
Argonne, Chicago

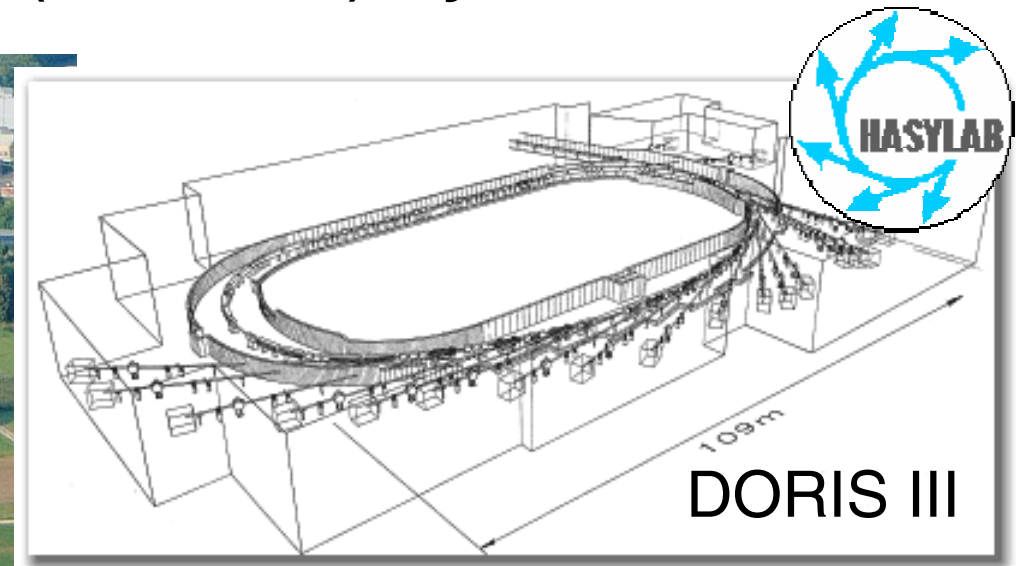
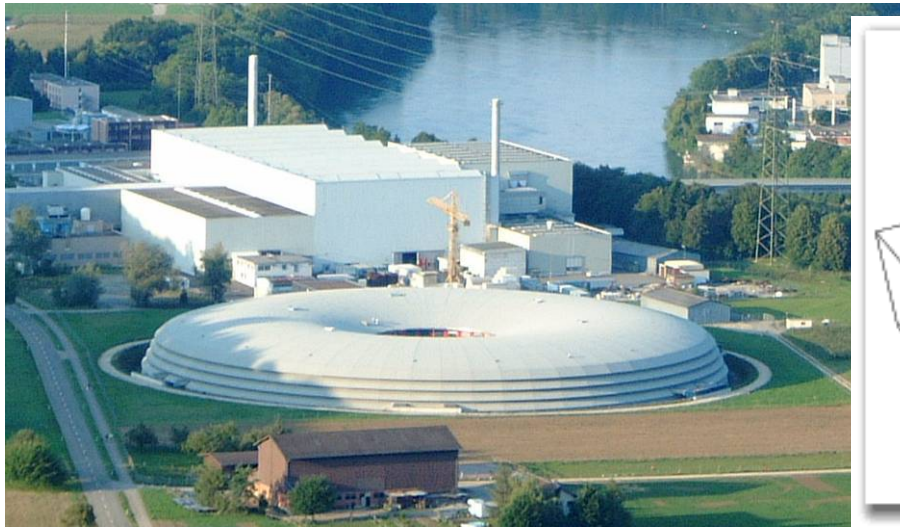




European Synchrotron Radiation Facility
Grenoble (France): $E = 6 \text{ GeV}$, $U = 844 \text{ m}$



Examples of smaller (national) synchrotrons



DORIS III



(Northern) German project

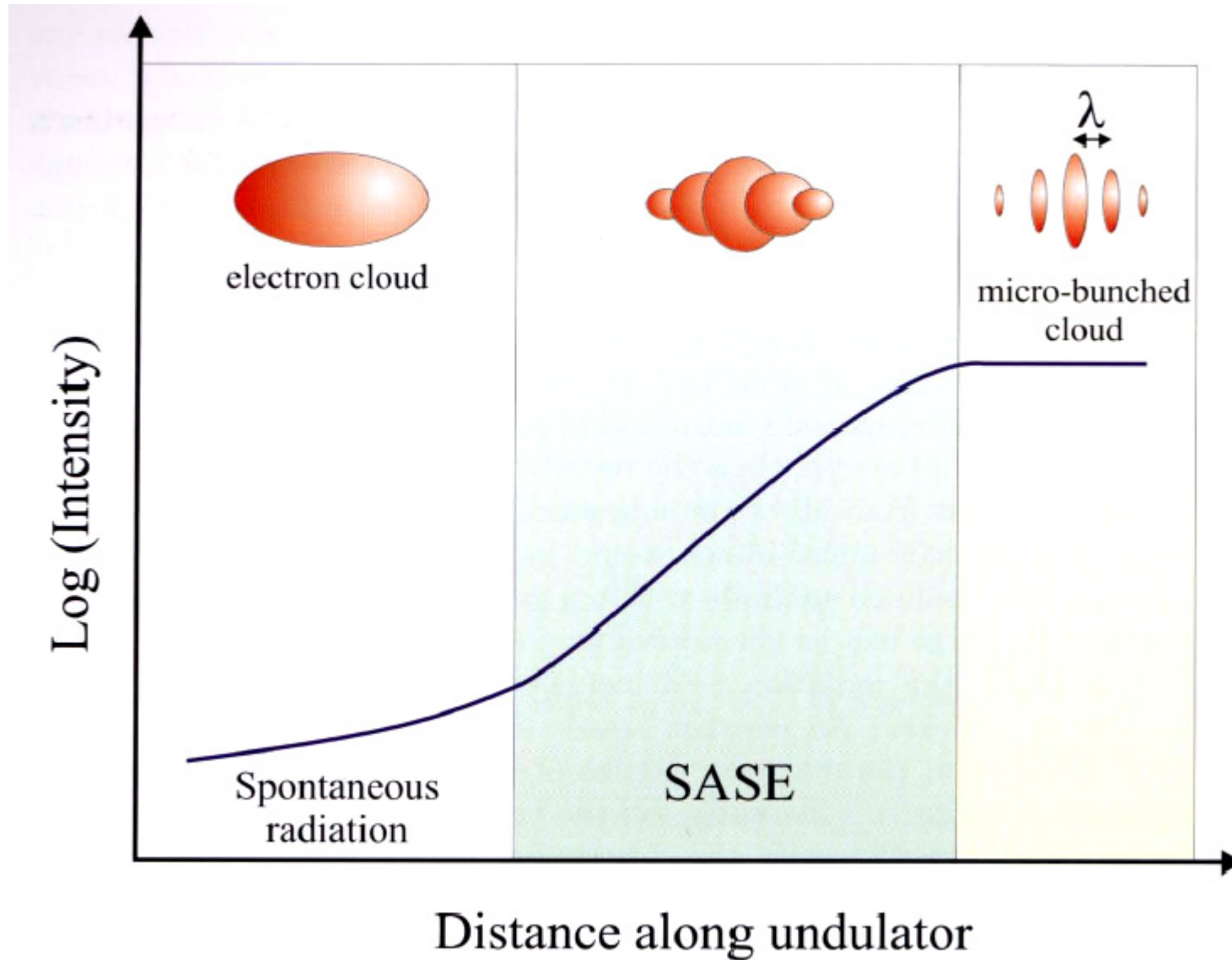


PETRA III, HASLAB, Hamburg

$E = 6 \text{ GeV}$, $U = 2304 \text{ m}$



Freier Elektronenlaser: Kohärenz



<http://xfelinfo.desy.de/de/artikel.fel-prinzip/2/index.html>